

Galileo Galilei (1564 - 1642)





	Axis of Rotation
	Invention of the telescope in 1610 (Galileo Galilei, Johannes Fabricius, Christoph Scheiner, and Thomas Harriot)
	Sunspots as tracers to measure differential rotation
	Rotation elements: (1) $i = 7.25^{\circ}$ angle of inclination between the ecliptic plane and the equatorial plane of the Sun and (2) Ω the ecliptic longitude of the ascending node of the Sun's equator
	Precession of the Earth's axis
	Heliographic coordinates
	Solar latitude ψ is the angular distance from the equator
	Colatitude: $\theta = n/2 - \psi$ (often used in theoretical work)
	Carrington rotations: intervals of 27.27 days starting 9 November 1853 ($\varphi = 0$ at the beginning of a new rotation)
	Conversion to ecliptic ccordinates
	$\sin \psi = \cos i \sin \beta - \sin i \cos \beta \sin (\lambda - \Omega)$
	$\tan \phi = \cos i \tan (\lambda - \Omega) + \sin i \tan \beta \sec (\lambda - \Omega)$
Nov	vember 3, 2005 Center for Solar-Terrestrial Research